

Using Swarm Satellite Data to Improve Global Geomagnetic Reference Modelling

Ciarán Beggan

British Geological Survey, Edinburgh, UK

With thanks to Susan Macmillan and William Brown



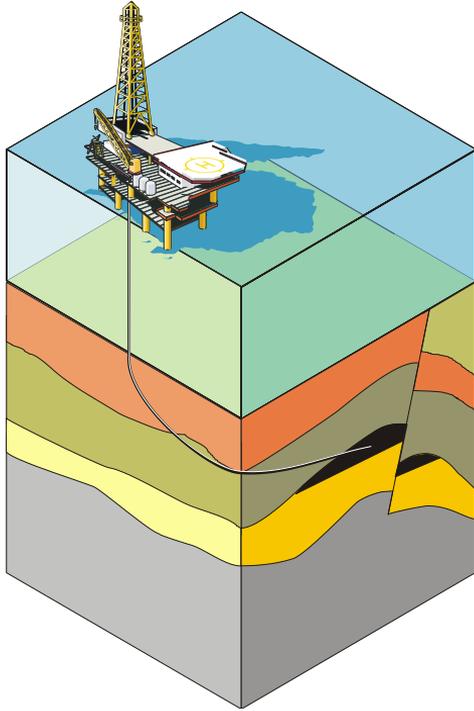
Speaker Information

- Dr. Ciaran Beggan
- Using Swarm Satellite Data to Improve Global Geomagnetic Reference Modelling
- 22-Sep-2016
- British Geological Survey

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BGS Geomagnetic Referencing



- BGM – global field model
 - includes quiet night time external field and long wavelength crustal field (> 300 km)
- IFR - combines internal field with estimates of local crustal field
 - higher spatial resolution of crustal field at drilling site (~1km)
- IIFR - combines IFR with estimates of field from external sources
 - higher time resolution (1 min)
 - supplied in real-time

Reducing uncertainties

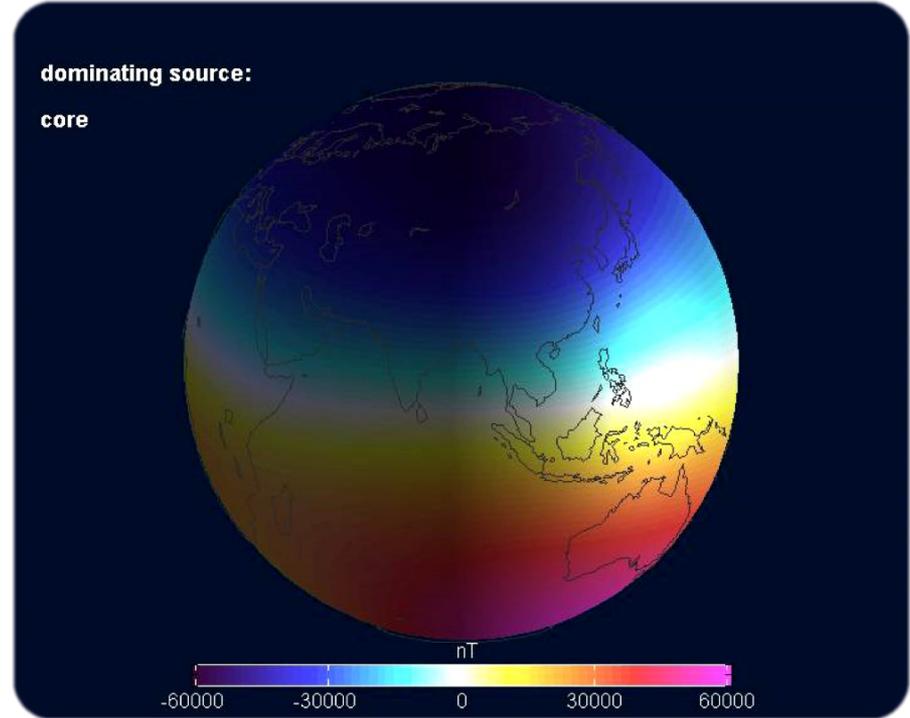
Overview

- The Earth's Magnetic Environment
- ESA Swarm Mission
- Secular variation: Jerks, IGRF and Model Updates
- Modelling Uncertainties
- Summary



Earth's Magnetic Field

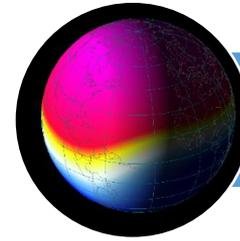
- Dominant **core field** varies over **months to years** ($\pm 60,000$ nT)
- Fields due to complex current systems in the **ionosphere** and **magnetosphere** vary from **seconds to years** (± 60 nT)
- Localised **crustal field** stable **through time** (± 10 nT)
- Now resolving **ocean tides** induced fields (± 2 nT)



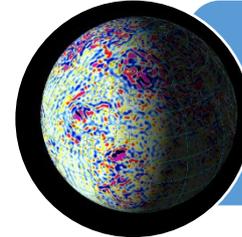
Credit: GFZ,DTU

Global Magnetic Field Modelling

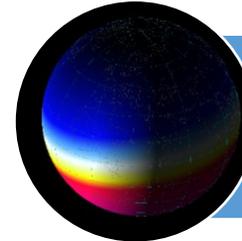
- Spherical harmonic 'degree' of model represents spatial resolution
- Produced annually since 1990s
- A high temporal and spatial resolution
- Most models describe time-varying main field and include crustal and some external fields
- Accurately quantified errors (compared against ground measurements)



Main Field



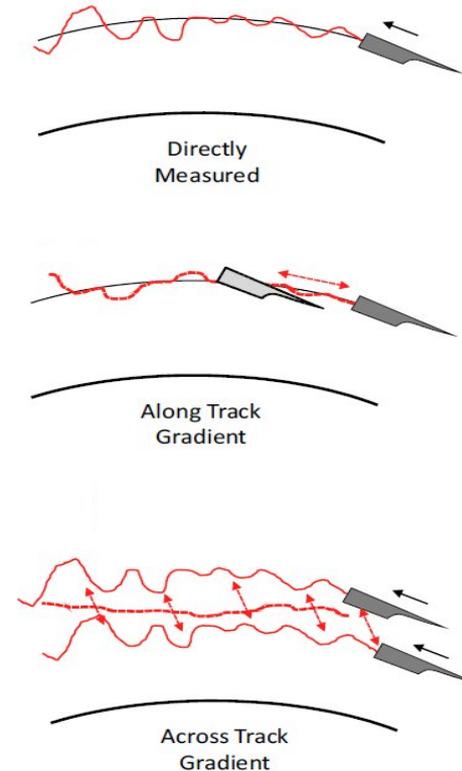
Crustal Field



External Fields

ESA Swarm Mission

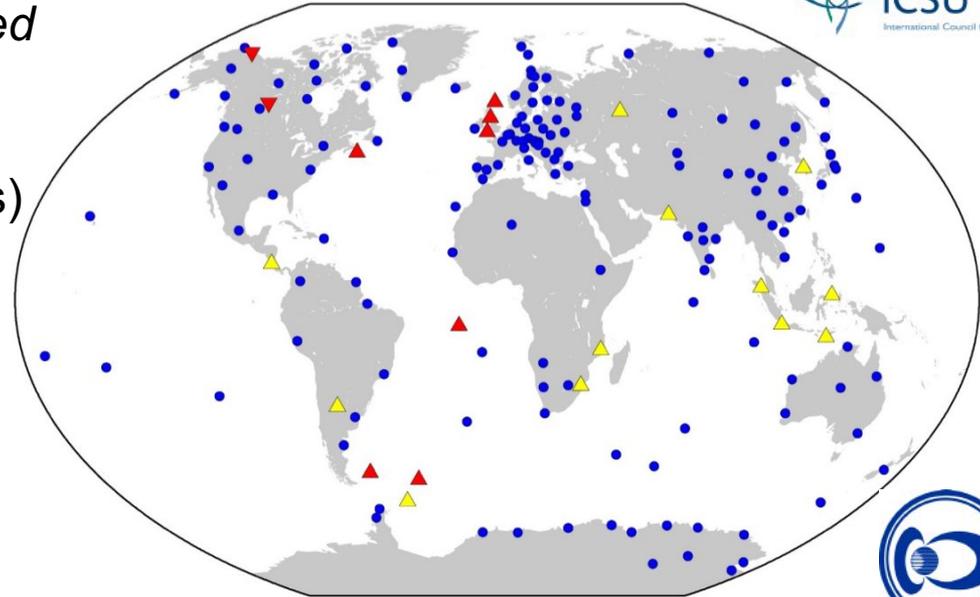
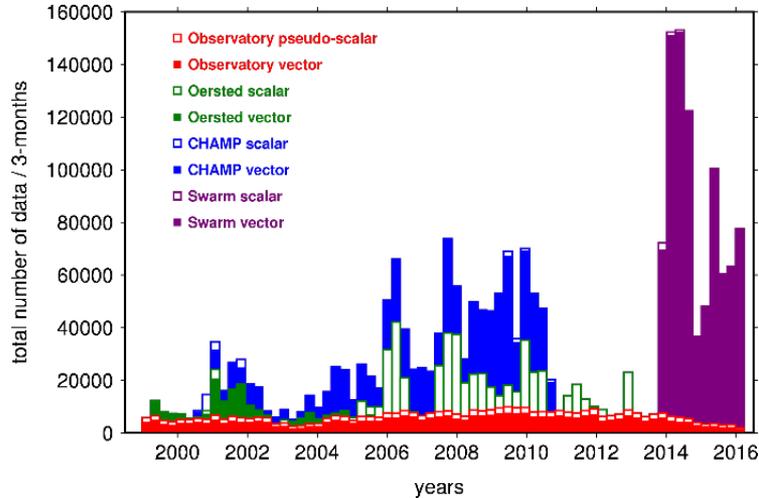
- CHAMP, Ørsted satellites have driven previous decade of modelling development
- Novel 3-satellite constellation
- Launched November 2013
- Two lower satellites (A and C)
- Swarm B flies in different local time orbit
- Unique multi-satellite magnetic gradients (A and C)
- Currently flying at ~450/510 km altitude, discussion next week in Edinburgh as to orbital evolution for 2017-2025



Other Data – Ground Observatories



- Ground observatories offer *stable fixed* measurement sites
- Removes the *time-space ambiguity*
- Offer long-term datasets (> 150 years)



- Geomagnetic observatories
- BGS/Halliburton observatories
- BGS observatories
- BGS/RMIB-assisted INDIGO observatories

Over 100 are INTERMAGNET standard



Wellbore Positioning Technical Section

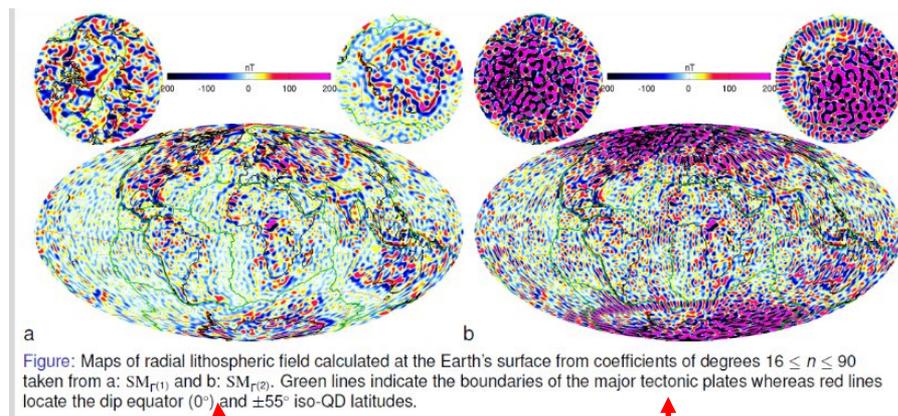


The Industry Steering Committee on Wellbore Survey Accuracy (ISCWSA)

Swarm offers great improvements

- Swarm offers new ways to use magnetic field data
- Gradients (various combinations)
- Boot-strapping crustal field
- Better external field descriptions
- Inclusion of smaller field components (e.g. tidal)
- Informs other research such as space weather effects on magnetic field and GPS etc

From Kotsiaros *et al.* (2016)



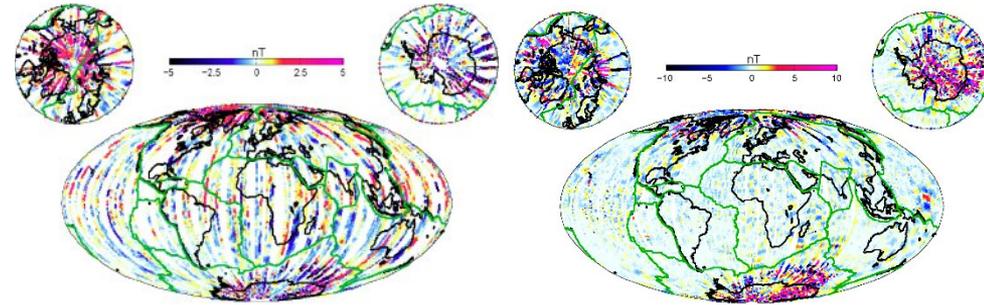
Model using radial and EW/NS gradients (better)

Model using differences of NS/EW gradients (poorer)

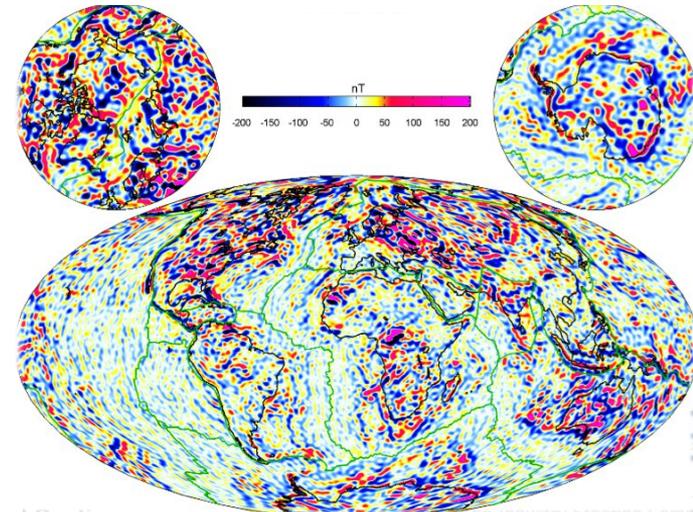
Gradients in use

- ‘Gradient’ data is calculated as the difference between two nearby measurements
- 15–50 second separation along track
- Nearest geomag latitude and time across track between A and C
- Sensitive to localised, small scale features
- Requires Swarm and CHAMP vector and scalar measurements (i.e. most of the data used)

Swarm cross-track and along-track radial gradients

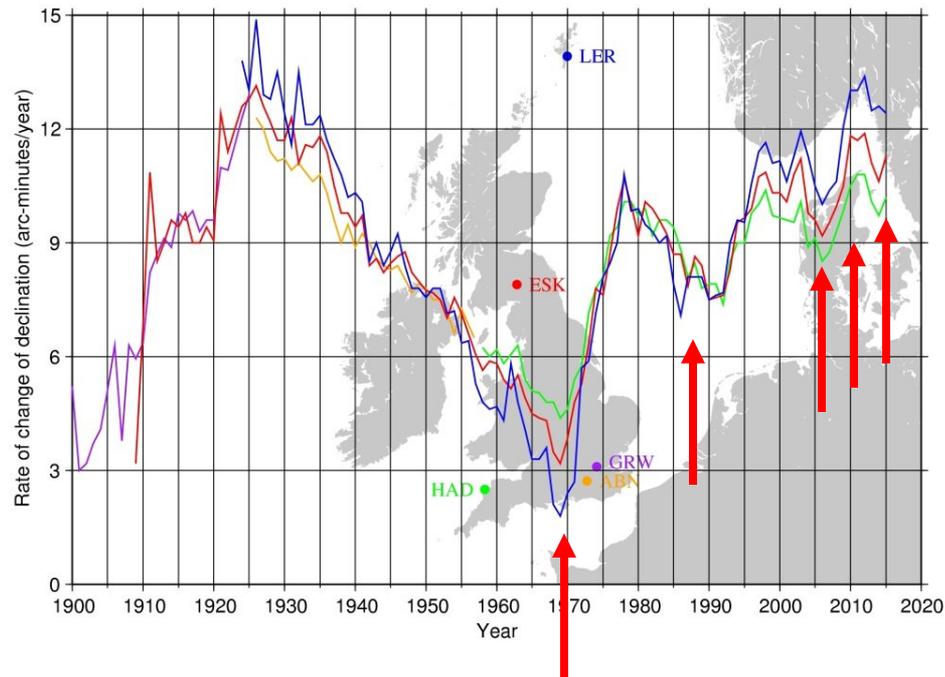


Final model: B_r at surface, $L = 16-133$



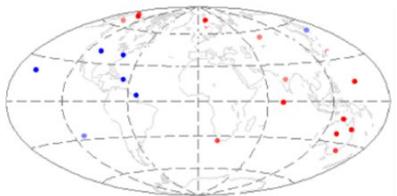
Secular variation of the main field

- Flow of liquid iron core generates secular variation at the surface
- Non-linear and constantly changing ('jerks')
- Current research looking to improve understanding

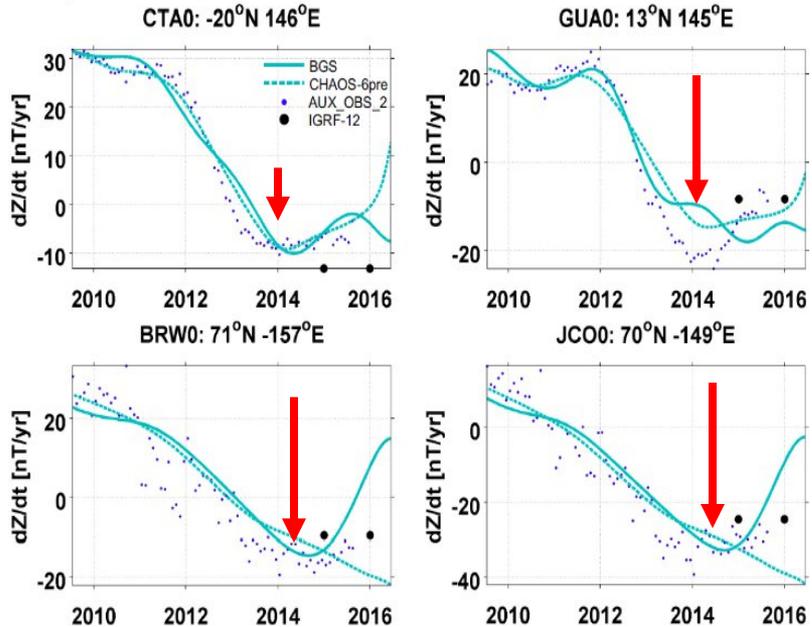


Imaging the 2014 Jerk

- Observatory distribution is sufficient to capture large scale SV of jerks but satellite data needed for fine detail



Measured jerks from obs. data

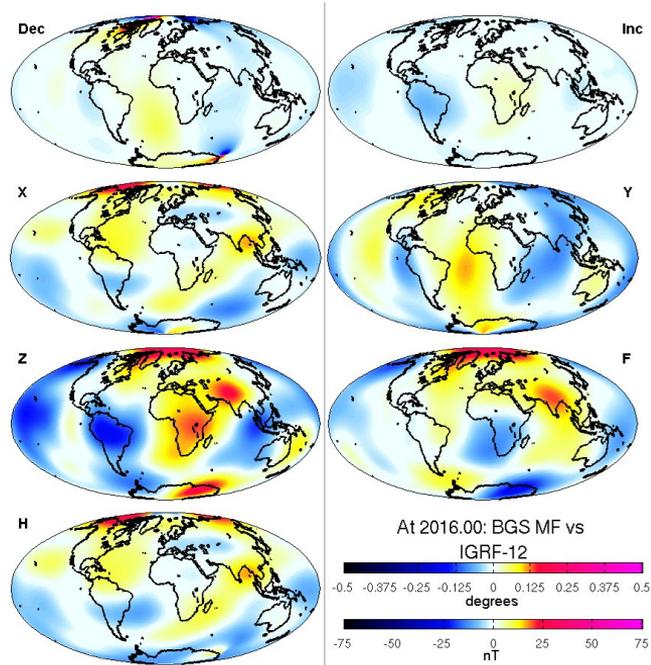


- Rapid Swarm and observatory data provision is key – the (BGS) model uses ~9 months of additional data to CHAOS-6 and IGRF-12



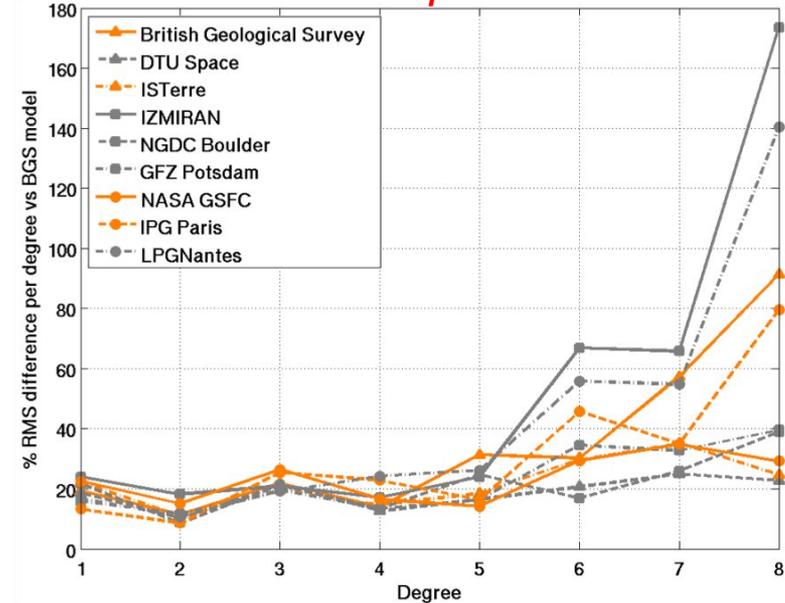
Jerks, IGRF and the importance of model updates

- Due to 2014 jerk, IGRF-12 prediction is different by **15.7 nT RMS** from recent core field model at 2016.0



44th General Meeting
September 22nd, 2016
Glasgow, Scotland, UK

IGRF-12 predictions



- Note, all 9 IGRF-12 candidate SV models **show 20% difference** from updated (BGS) model after one year into 5 year lifetime

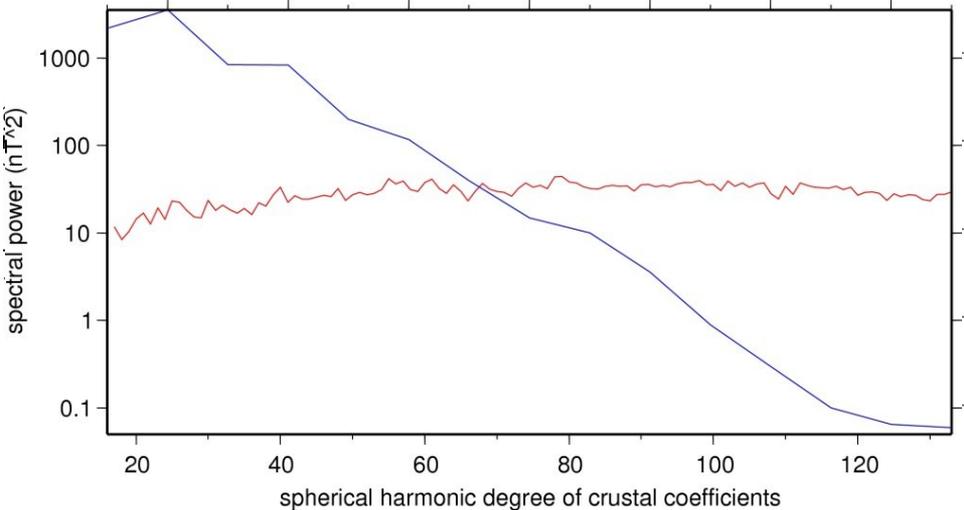


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Global mean power spectra



(following convention of logarithmic Y axis which emphasises small values)

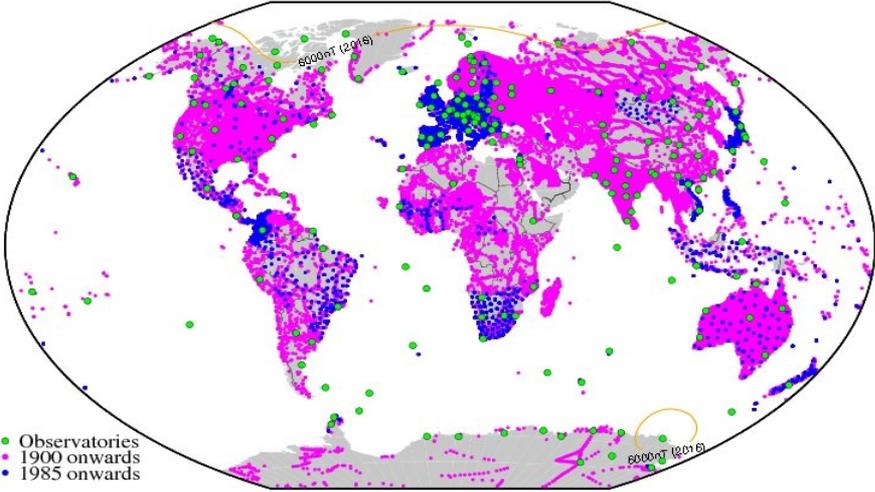
- Total power in core field (up to degree 15)
 $\approx 1,912,660,000 nT^2$
- Total power in core field change
 $\approx 7,750 (nT/yr)^2$ **every year**
- Total power in crustal field (degree 16-133)
 $\approx 3,410 nT^2$
- Degree 133 is about the current max with satellite data
= a *globally homogeneous* dataset
- Note: locally the crust can vary significantly

→ Accurately modelling the core field on a global scale arguably more important than crust

Update on model uncertainties

Original analysis 2008 (SPE119851)	Comparisons were made using 6-8k vector data and a degree 45 model
New analysis 2016	Comparisons were made in September 2016 using 8-10k vector data avoiding poles ($H < 6\mu T$) and a degree 133 model

Locations of observatory and magnetic survey stations with usable data



New model uncertainties

Confidence Level	Original Declination Limit (degrees)
68.3% (1 σ if Gaussian)	0.148
90%	0.419
95%	0.823
95.4% (2 σ if Gaussian)	0.874
99%	1.641
99.7% (3 σ if Gaussian)	2.613

- These errors (attributable to the local crustal field) to be combined with
 - estimates from hydrocarbon geology locations
 - external field
 - predictive core field errors
- Objective is one new error look-up table (scalable 1-sigma values) for annually revised high-degree global models, ideally with all-party agreement

(Using vector survey data 1985 and onwards only in both cases)

Locations of oil fields with local magnetic data



Summary

- Annually updated models necessary to counter large and unpredictable rapid changes from Earth's core
- Uncertainties are lowering but care needed not to misunderstand what global models can do
- Swarm is promptly delivering a large quantity of highly accurate measurements
- Swarm gradient data offers unique global resolution of small scale field, especially as orbit lowers



Acknowledgements & References

- Swarm: European Space Agency and Swarm Data, Innovation and Science Cluster
- Ground-based data: institutes worldwide with geomagnetism programmes, World Data Centre for Geomagnetism, survey companies Halliburton, Baker Hughes, Schlumberger, Weatherford
- SPE119851: Macmillan & Grindrod, 2010. Confidence Limits Associated With Values of the Earth's Magnetic Field used for Directional Drilling. *SPE Drilling & Completion*, 25(2), 230-238. doi: 10.2118/119851-PA
- Kotsiaros, S. (2016) Toward more complete magnetic gradiometry with the Swarm mission, *Earth Planet Space*, 68, 130. doi:10.1186/s40623-016-0498-x

